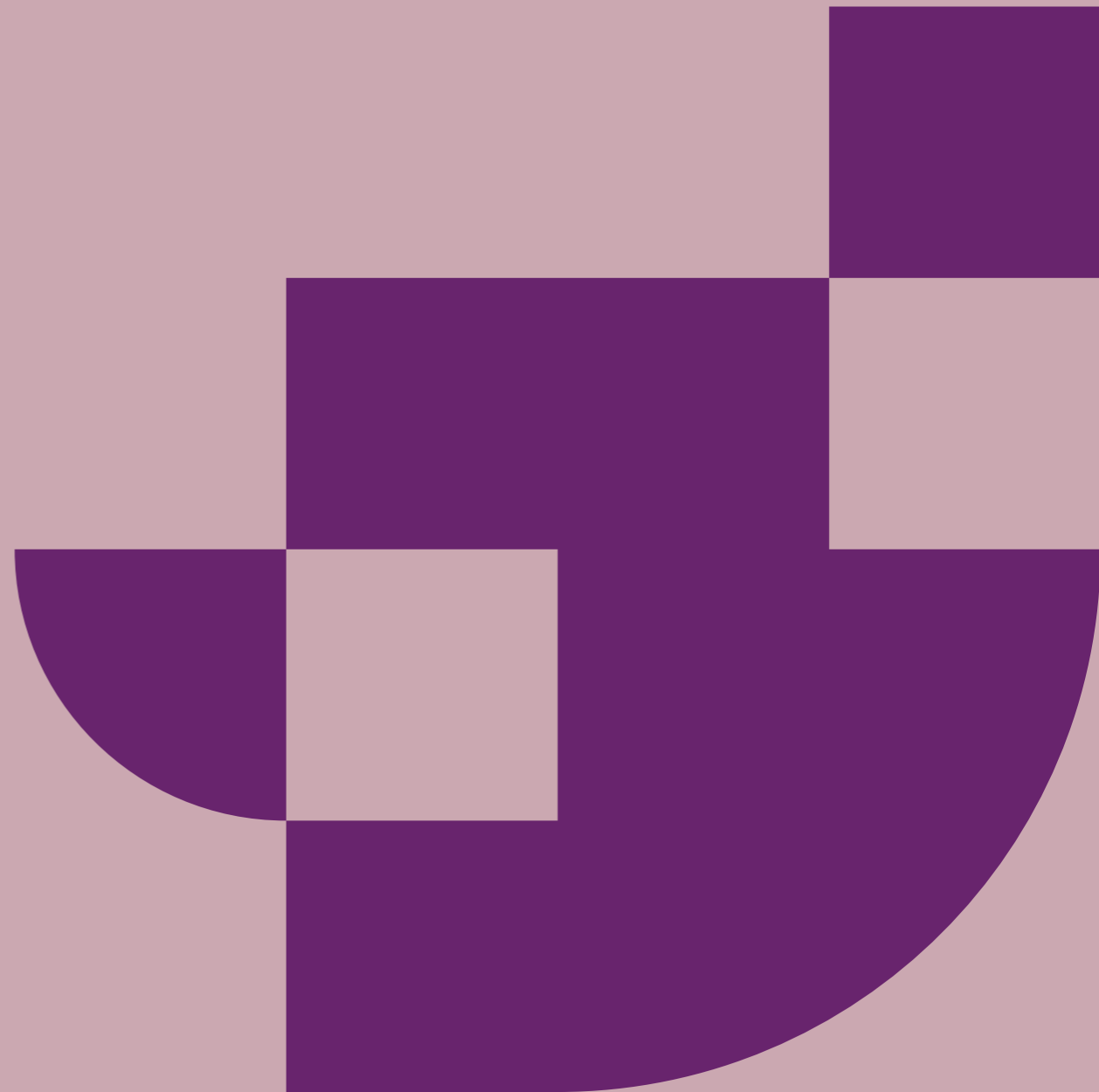


An outflow equilibrium model for the solar corona

Prof. Anthony Yeates
with Dr Oliver Rice



UK SWSE Mtg, Exeter, 10-Sep-2024

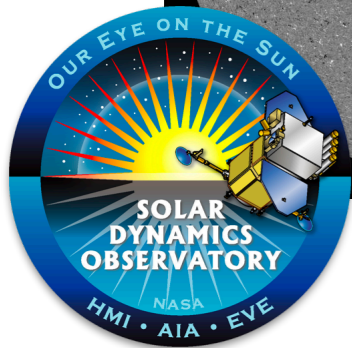
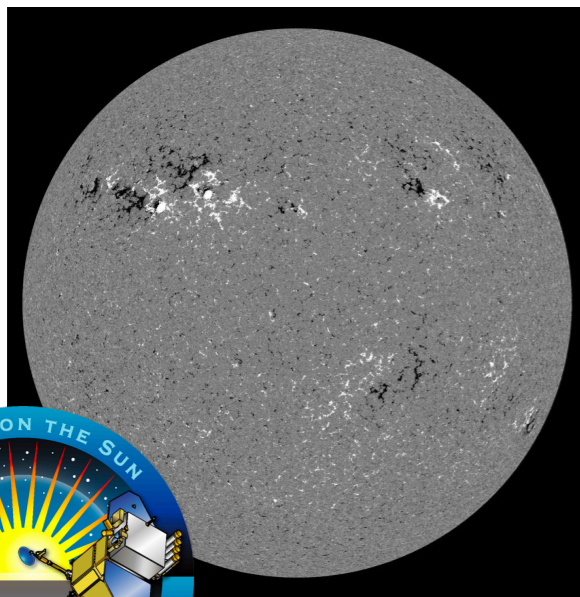


- ▶ **Magnetic models** are needed to extrapolate from solar surface observations to the inner heliospheric boundary.

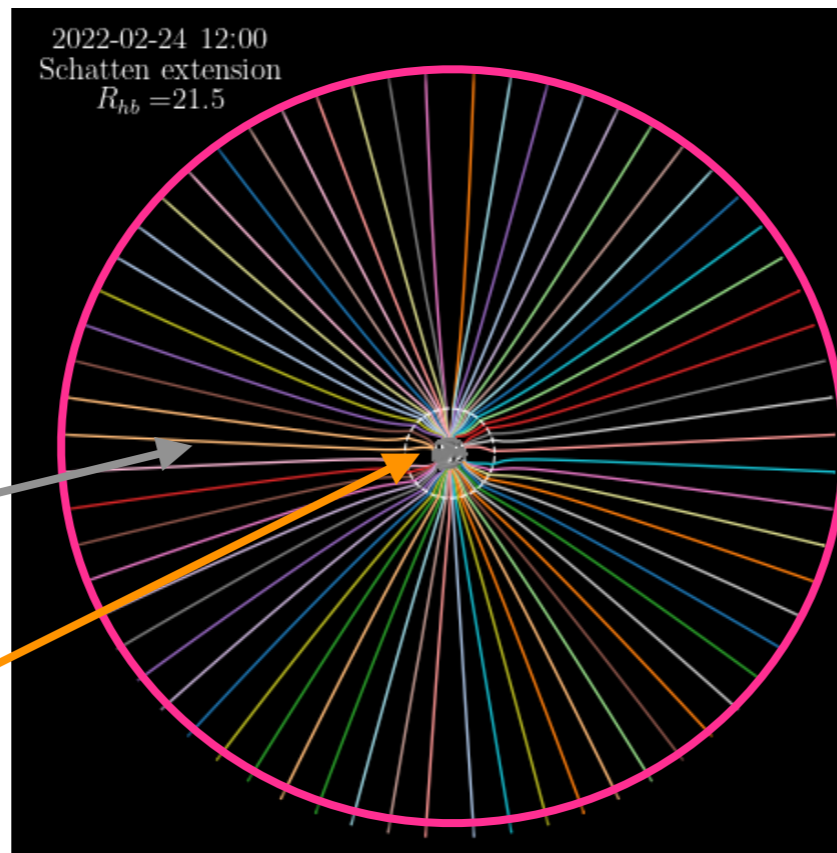
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e.g. SWEEP pipeline [SWIMMR/S4]

Satellite data



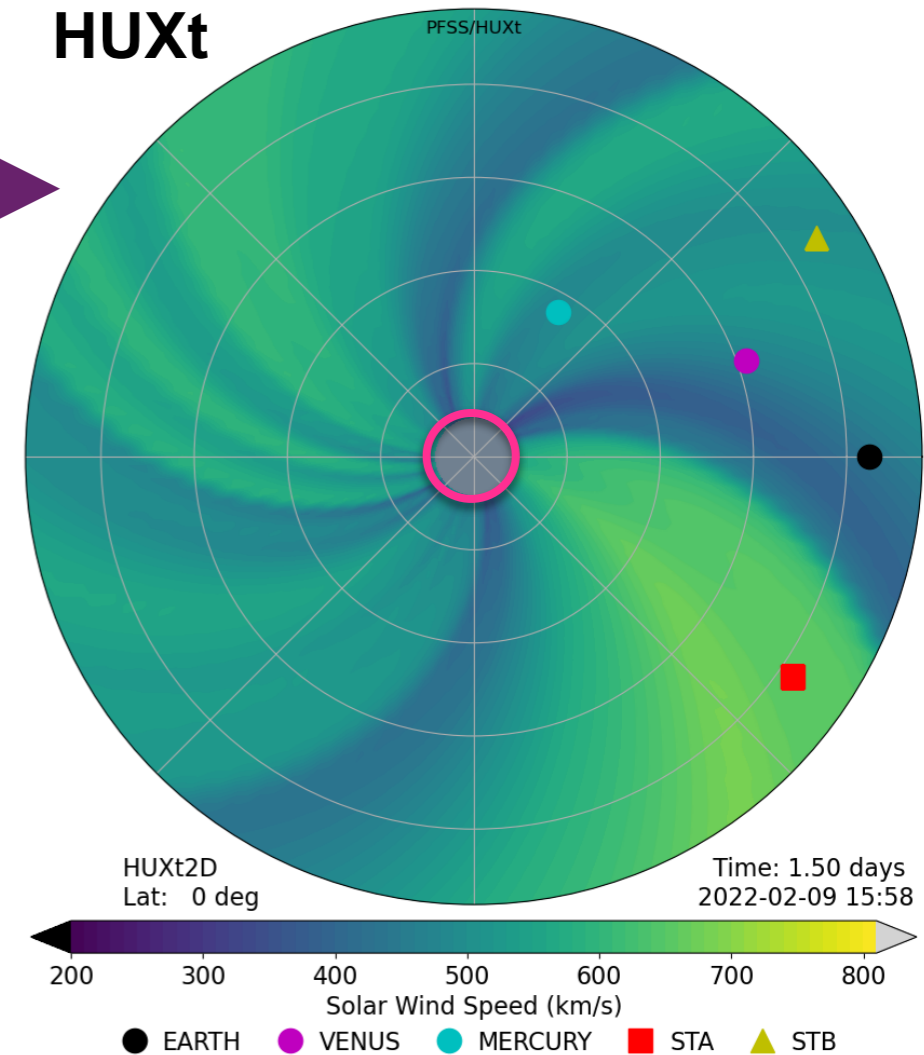
Coronal magnetic model



Schatten current sheet model

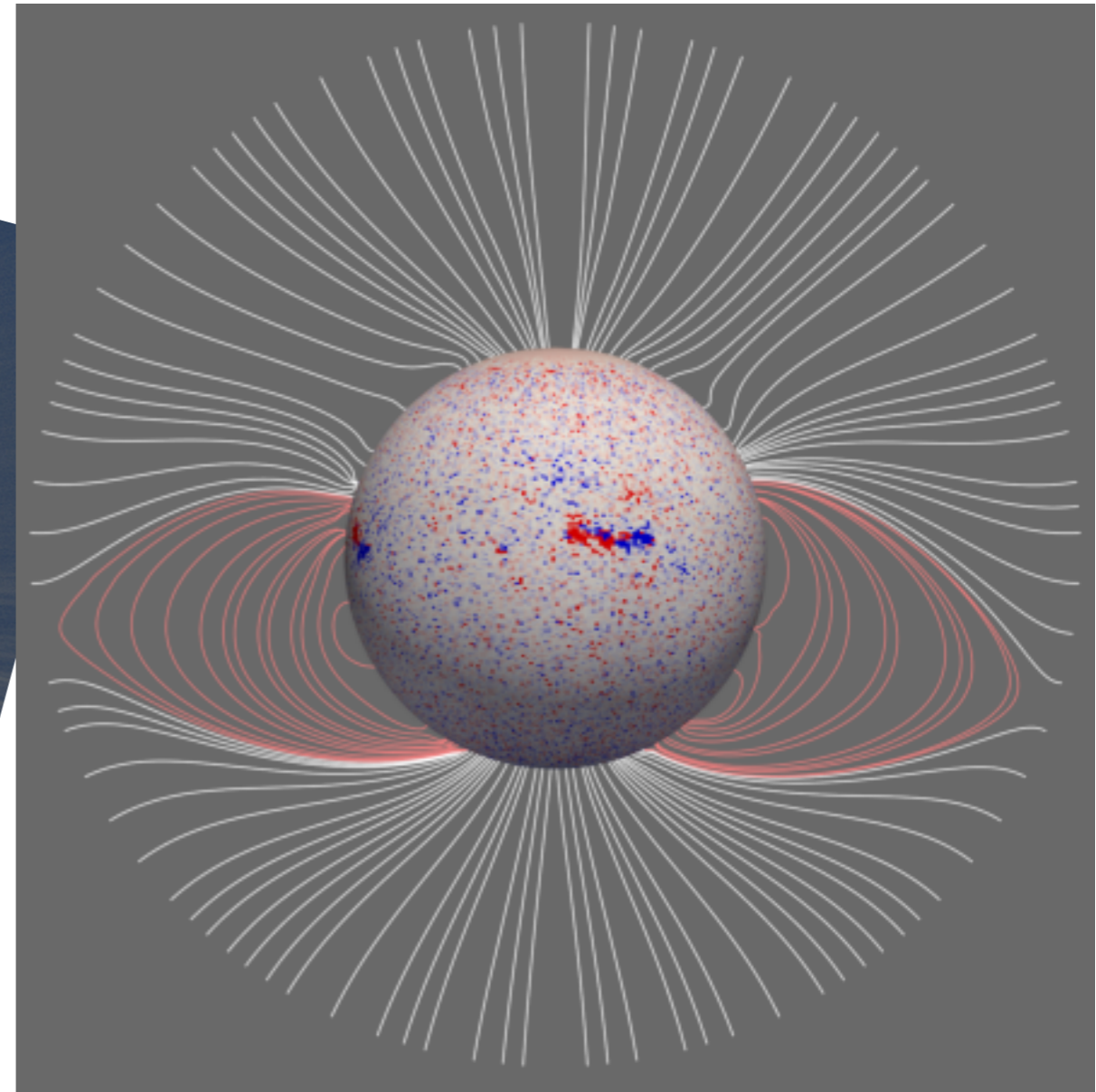
Main coronal model (up to $2.5R_{\odot}$)

HUXt



- ▶ Current operational forecasts use the **PFSS (Potential Field Source Surface)** model.

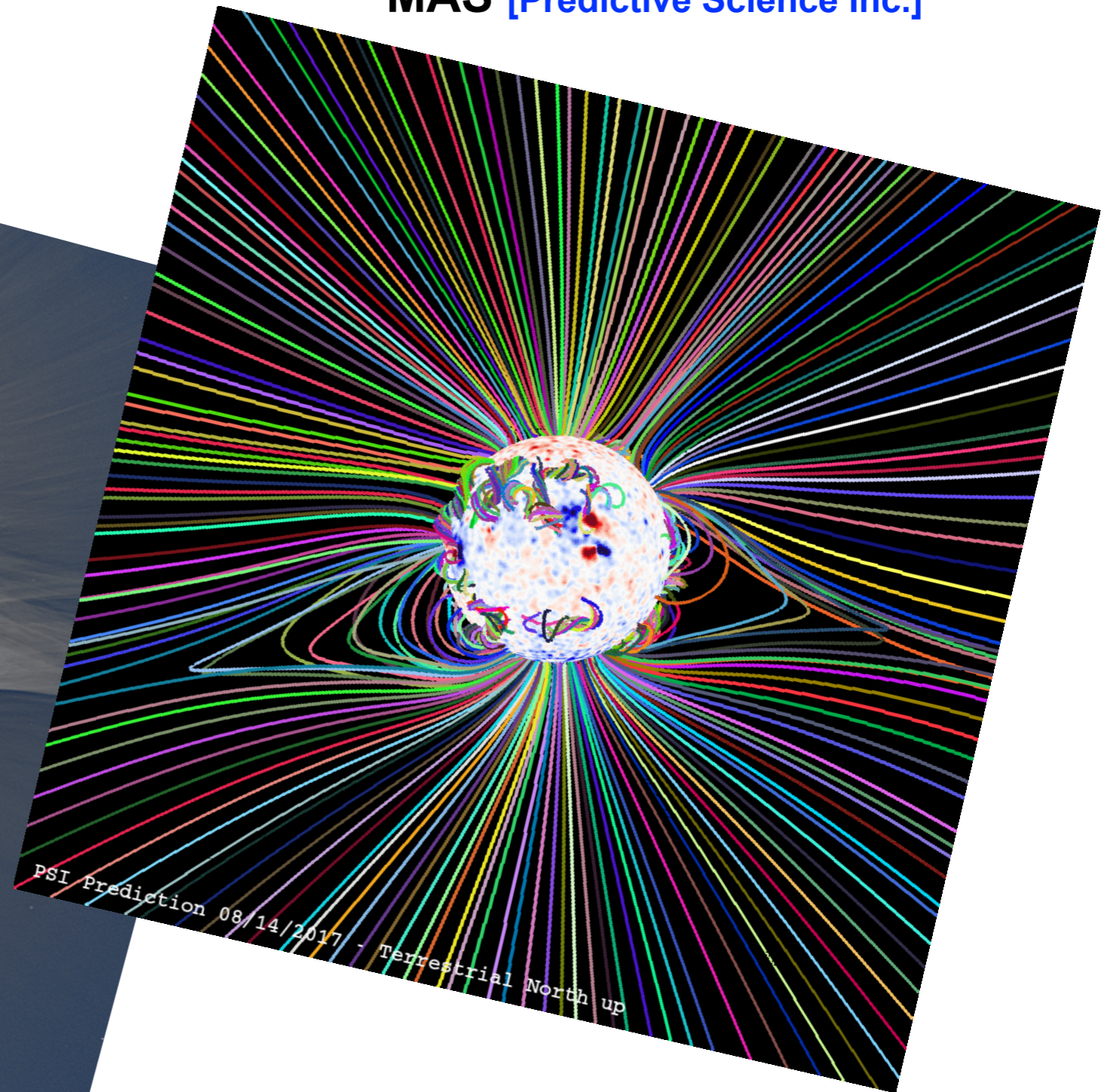
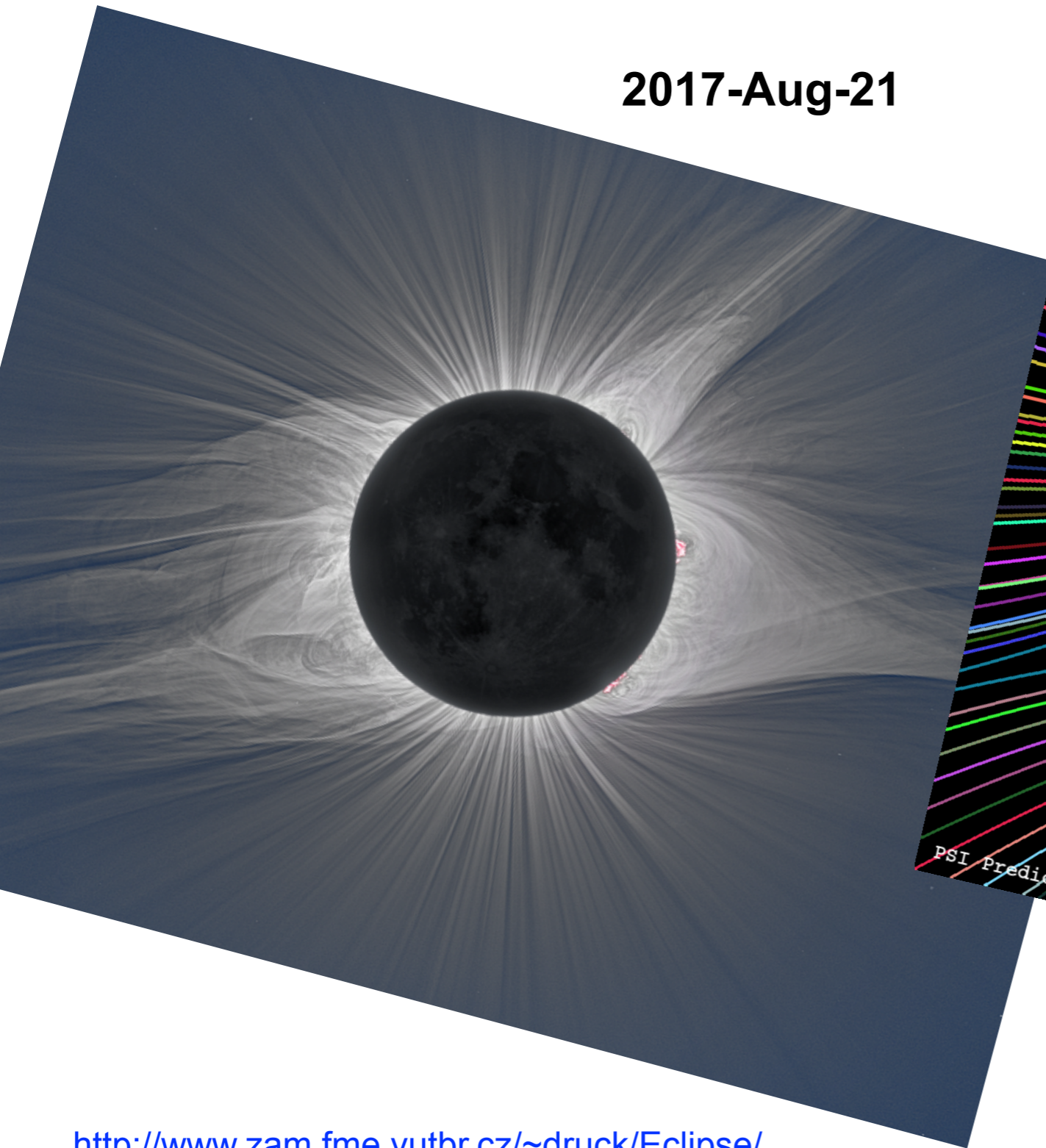
2017-Aug-21



► **cf.** full MHD modelling.

MAS [Predictive Science Inc.]

2017-Aug-21



<http://www.zam.fme.vutbr.cz/~druck/Eclipse/>

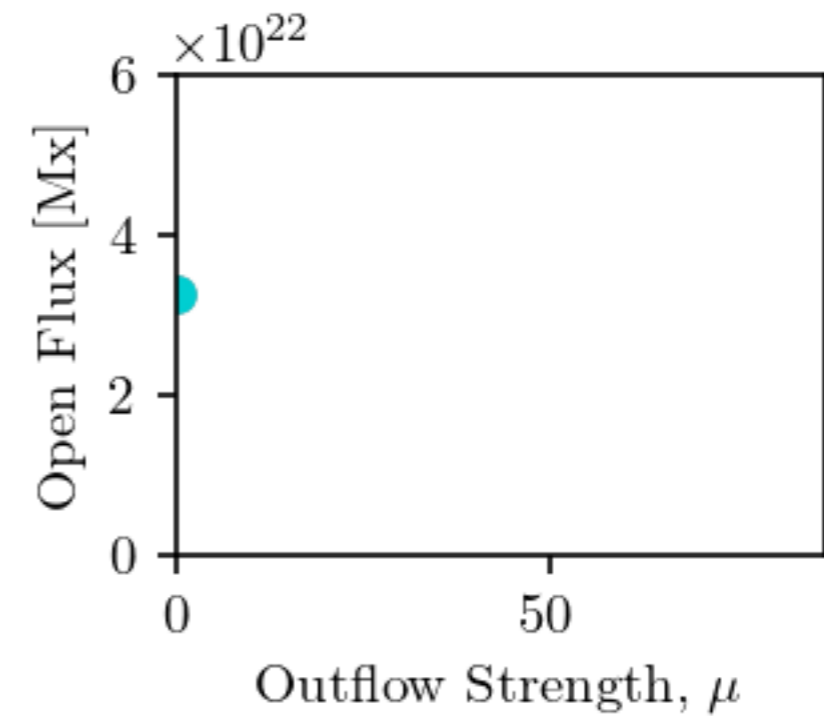
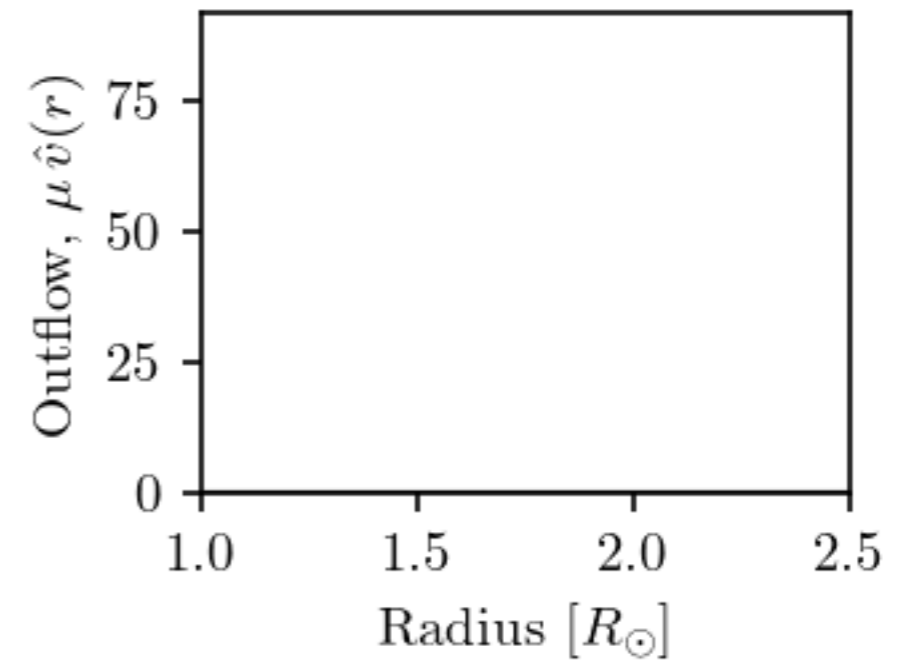
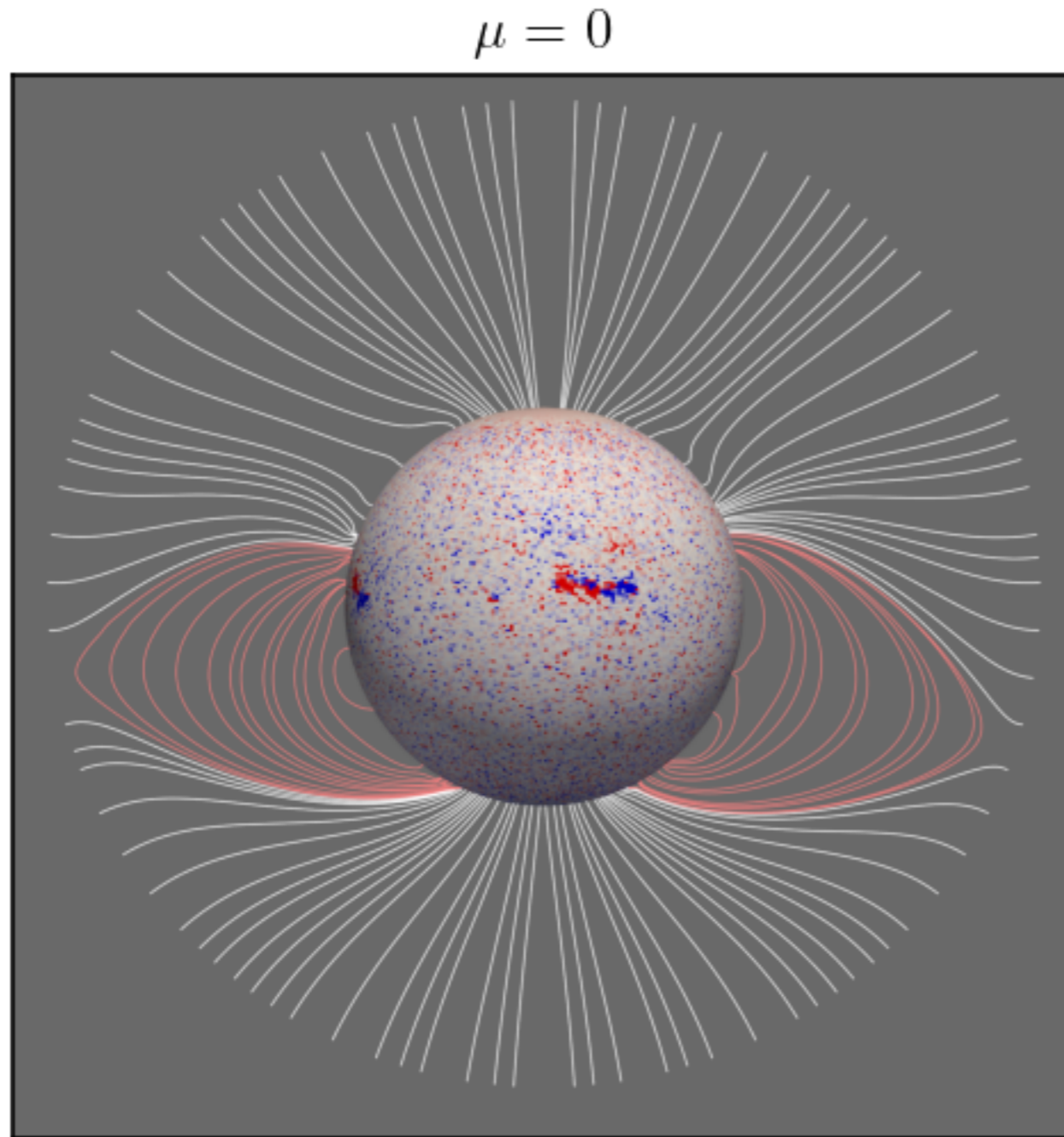
<https://www.predsci.com/corona/aug2017eclipse/>

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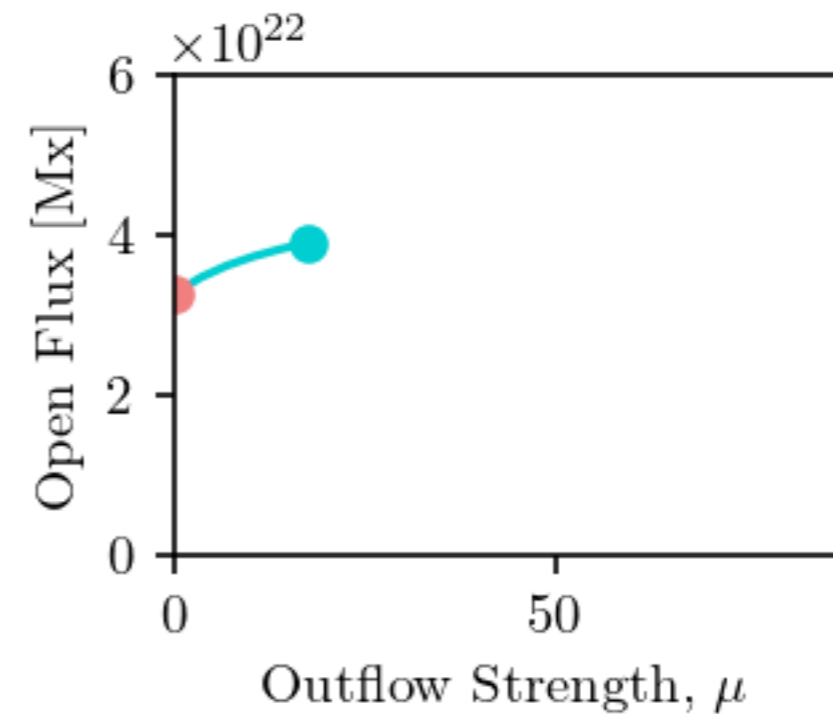
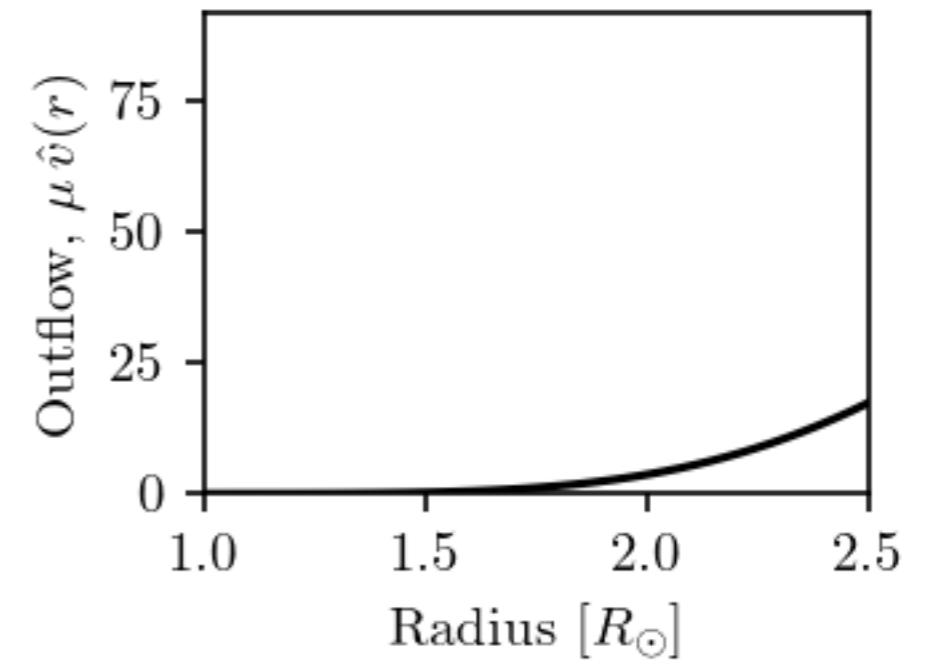
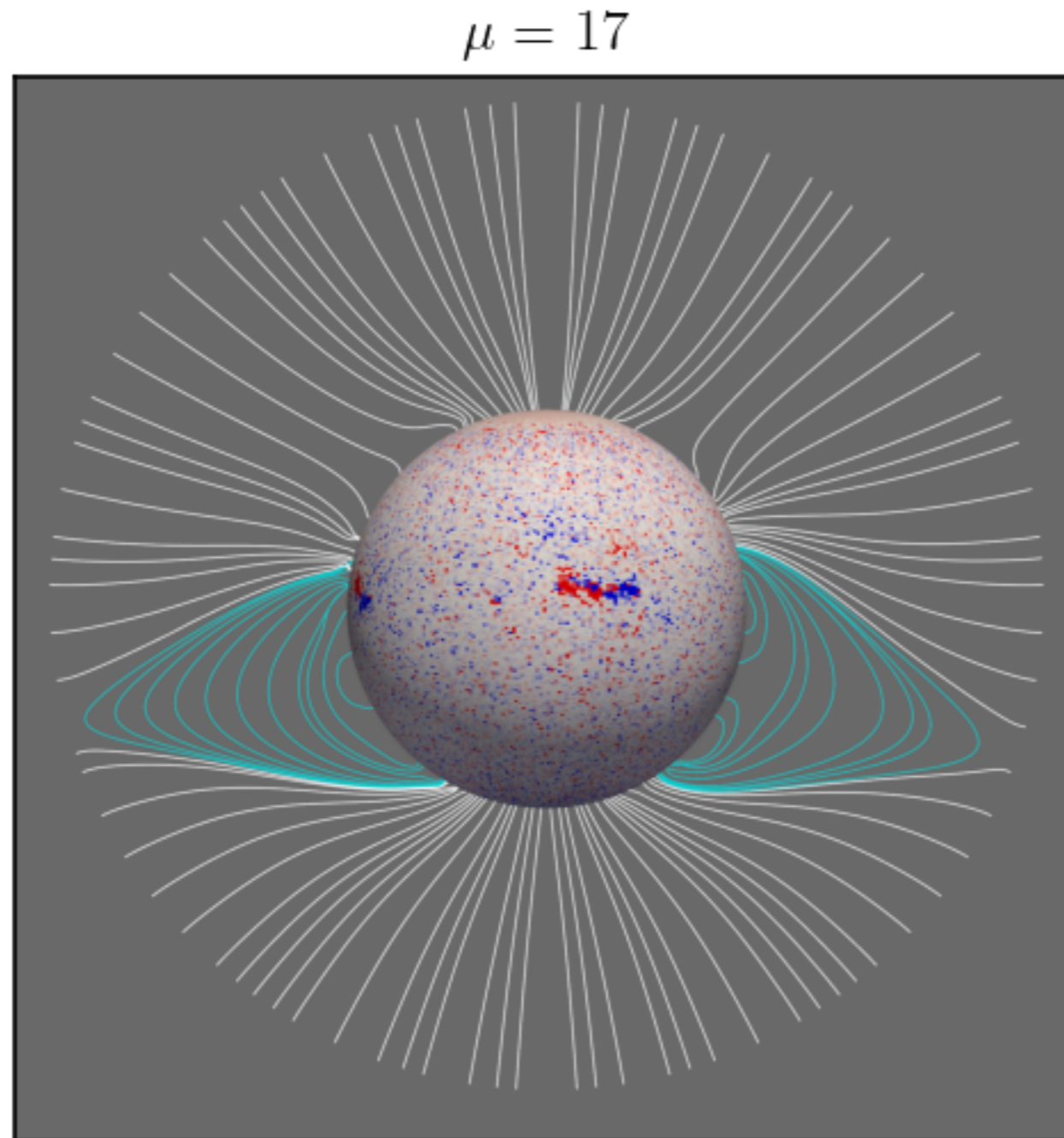
Rice & Yeates, *ApJ* **932**, 57 (2021)

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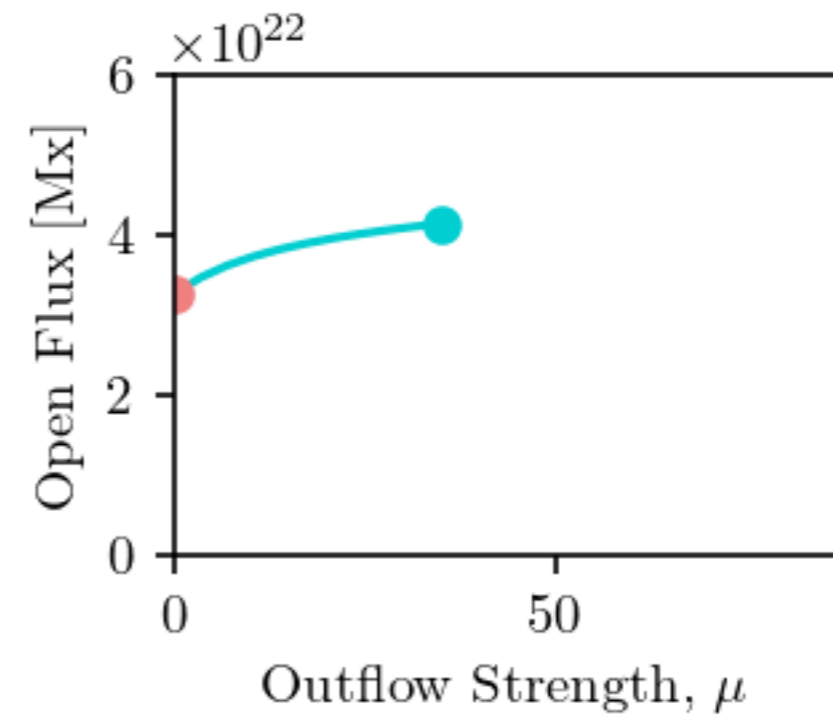
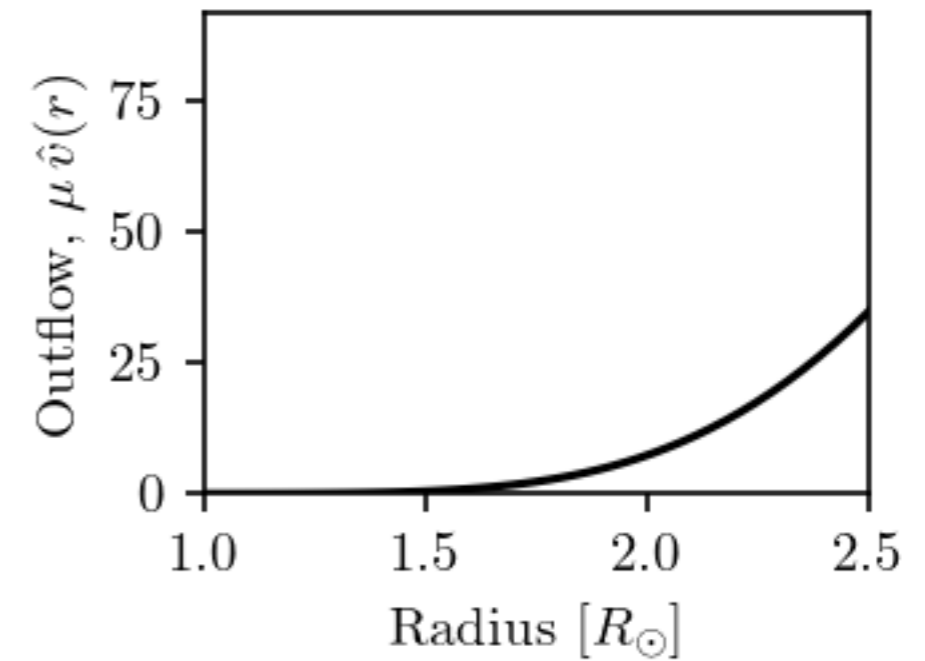
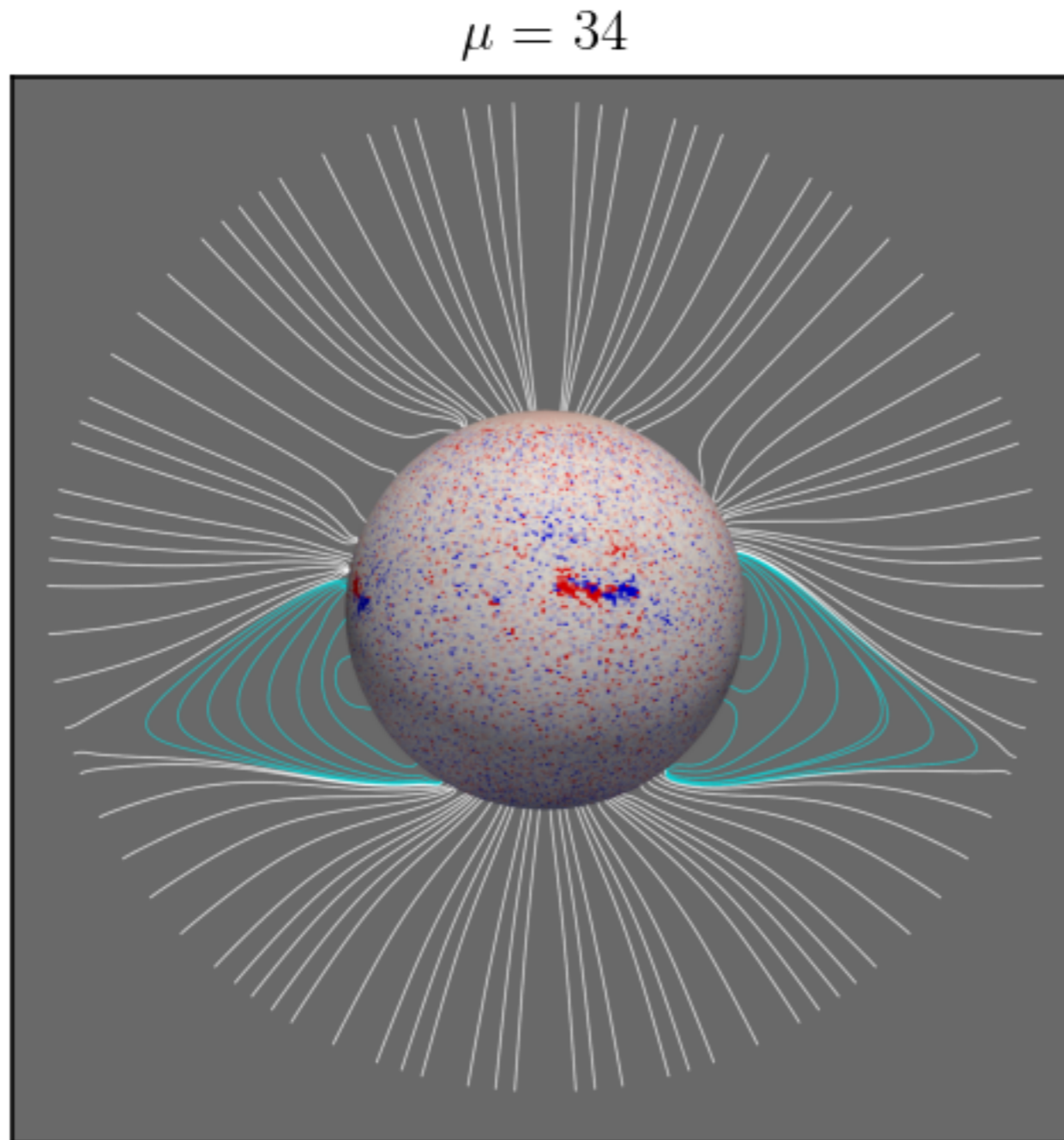
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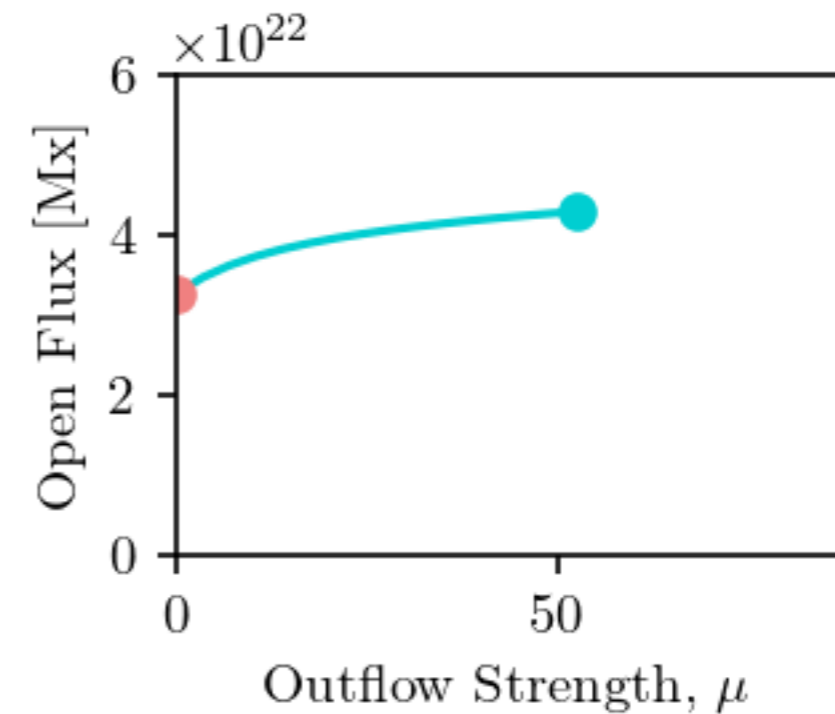
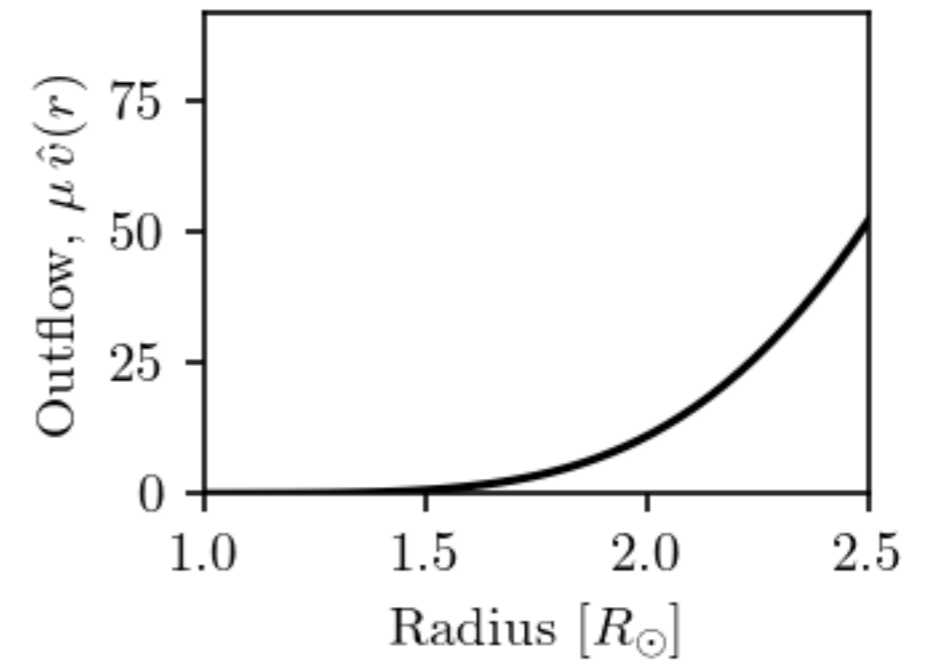
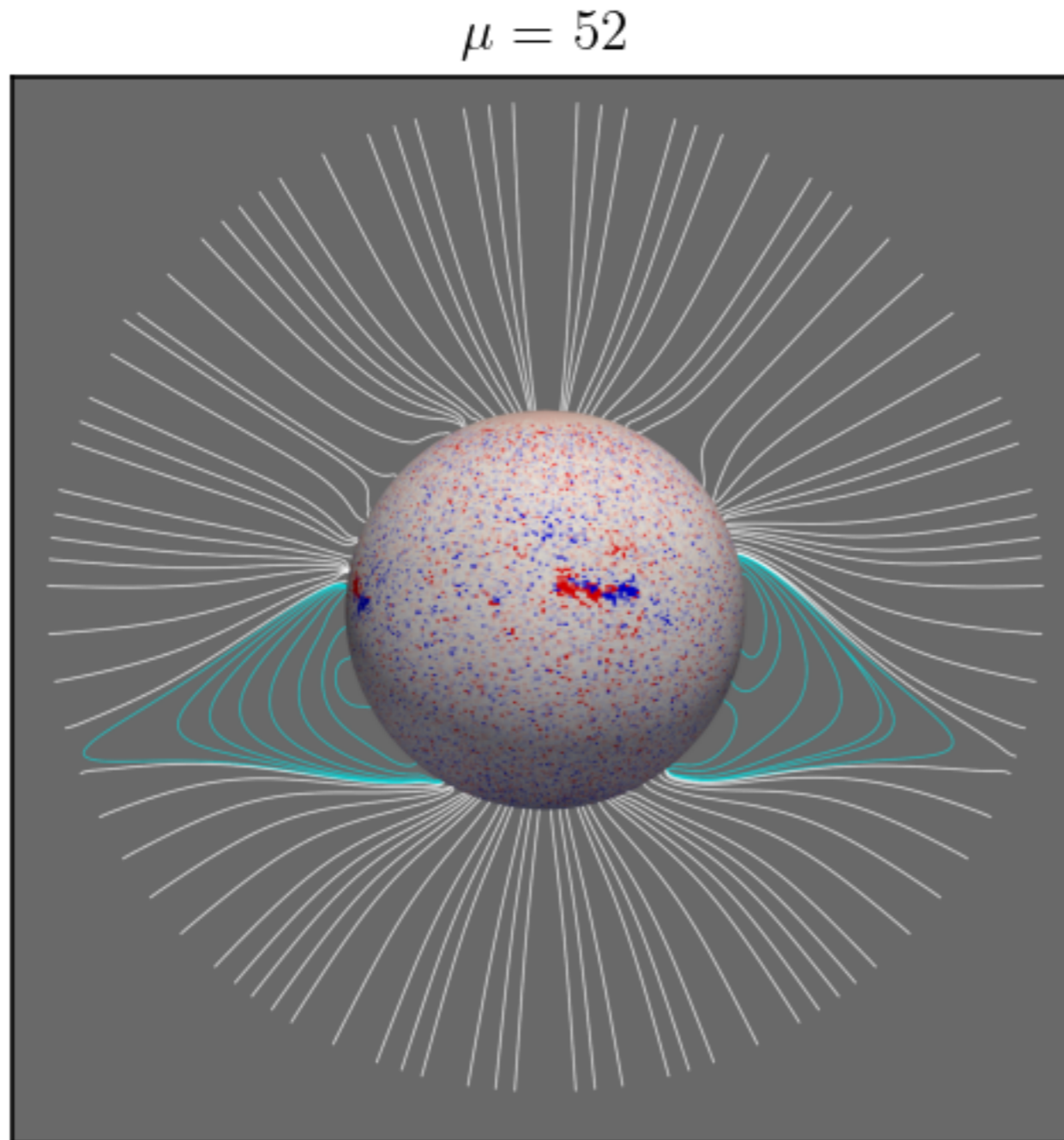
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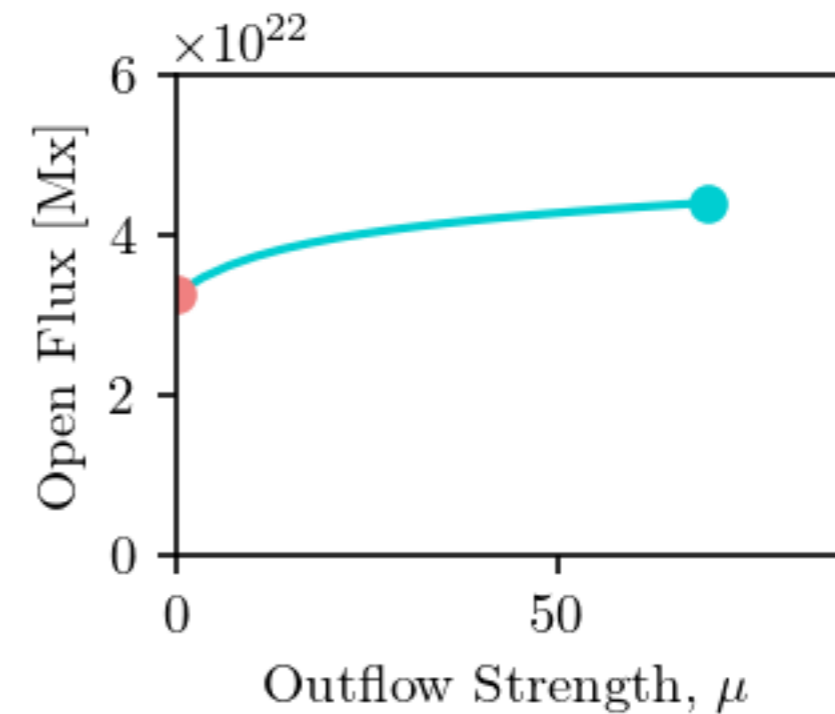
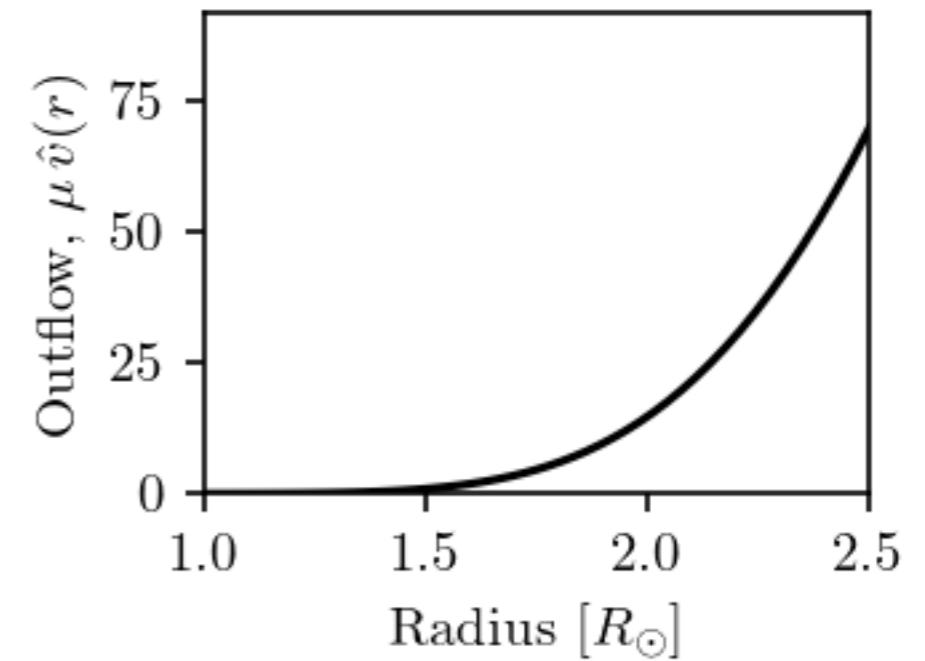
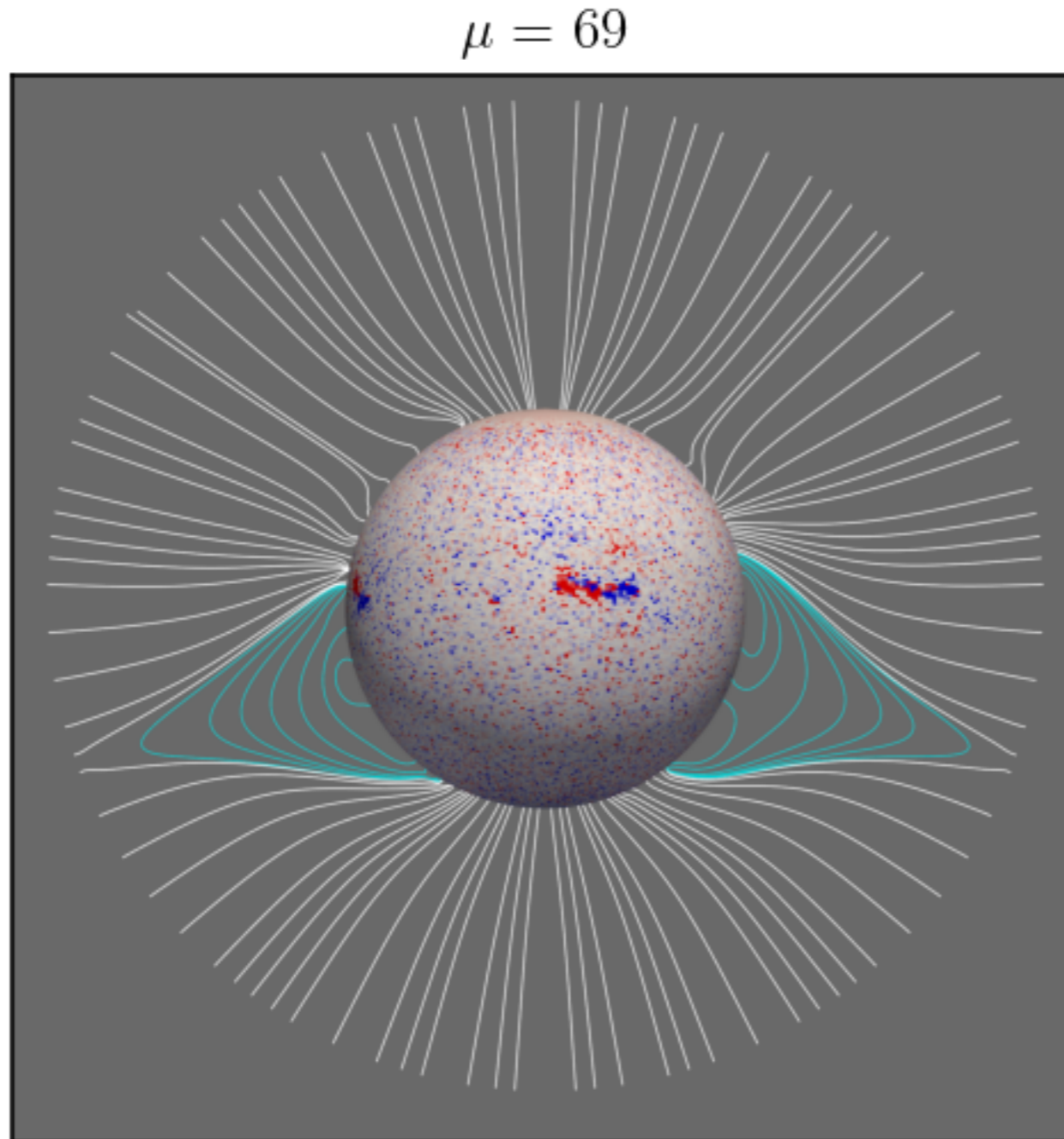
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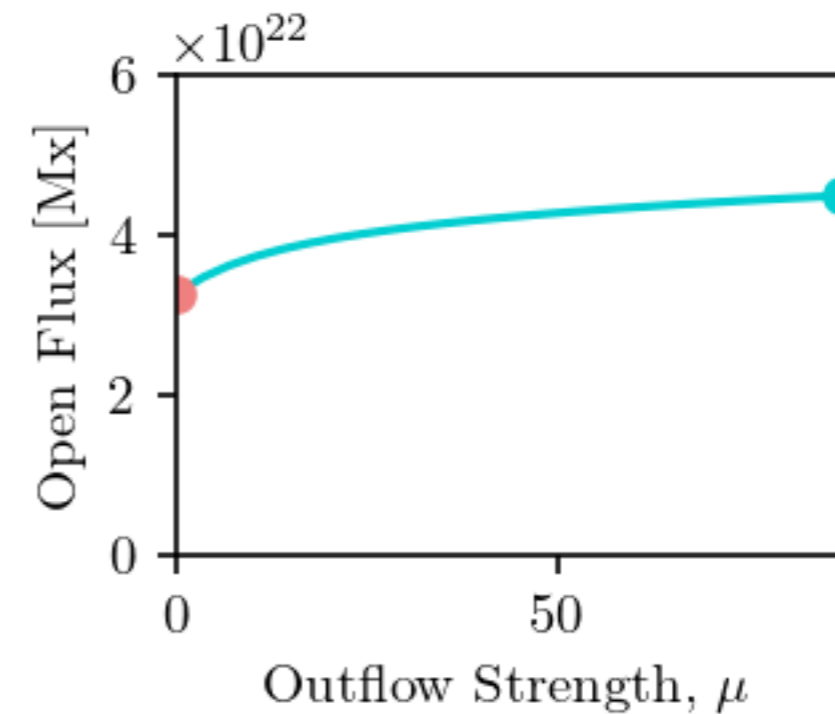
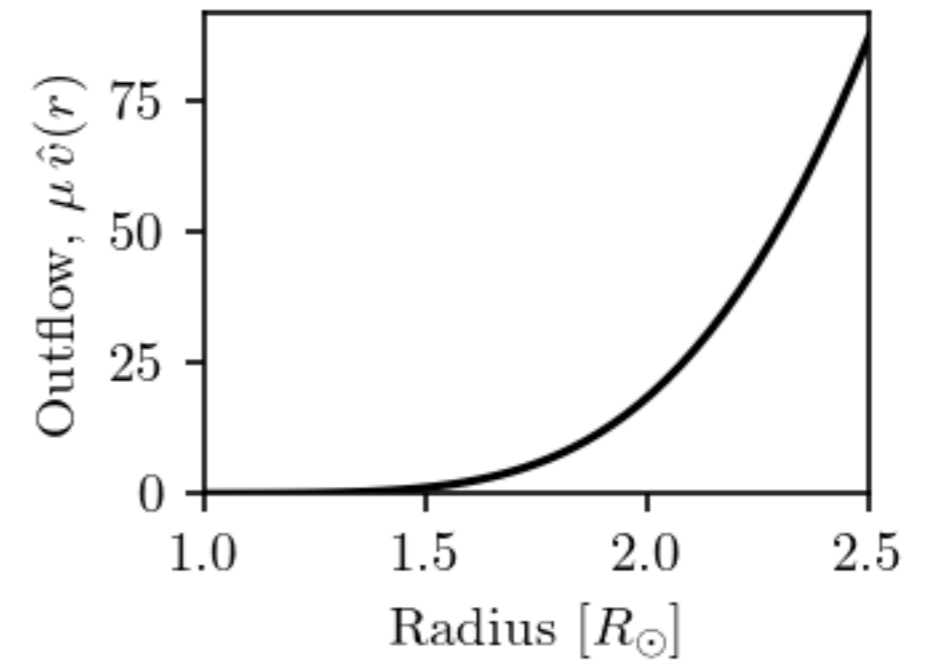
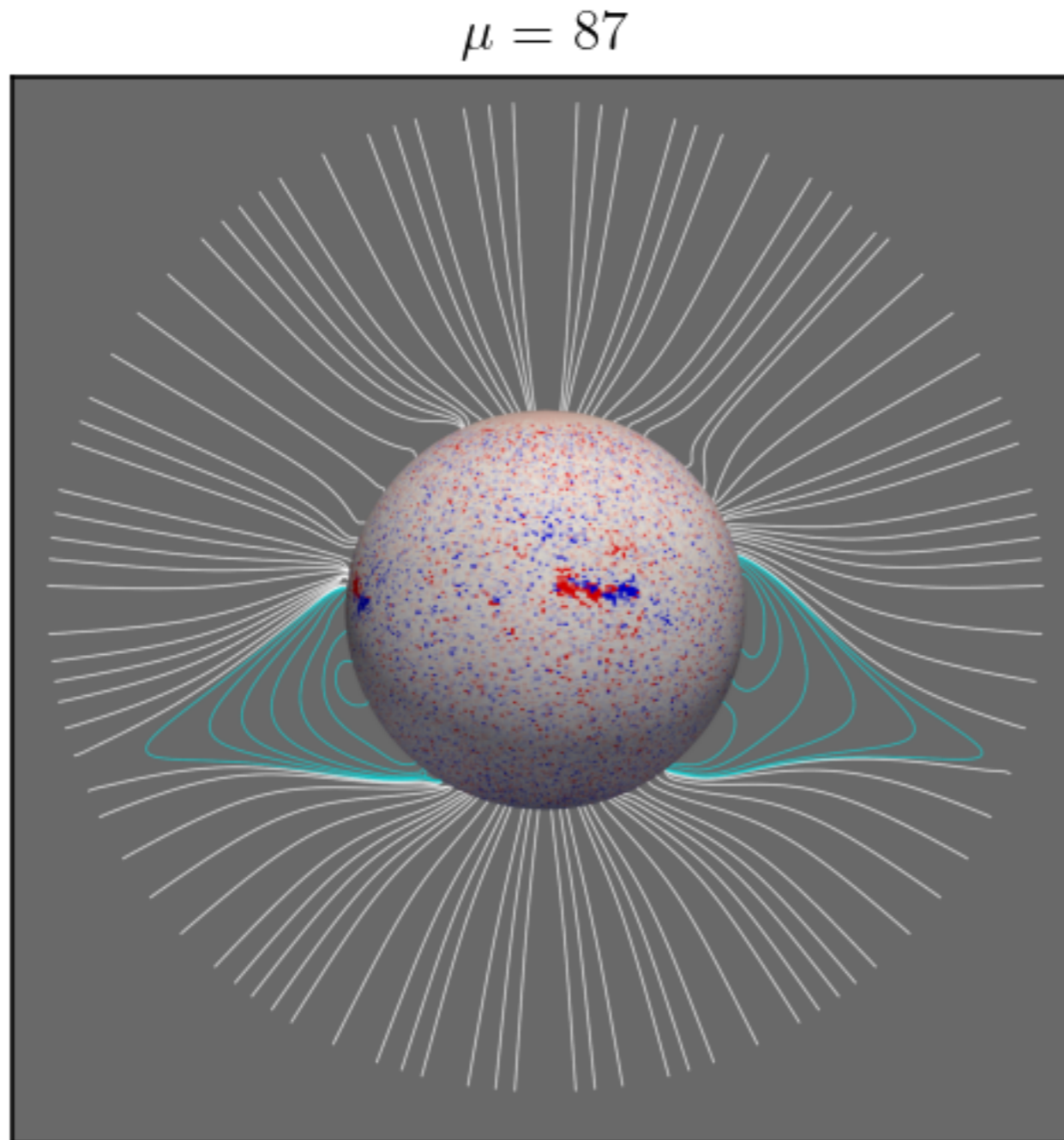
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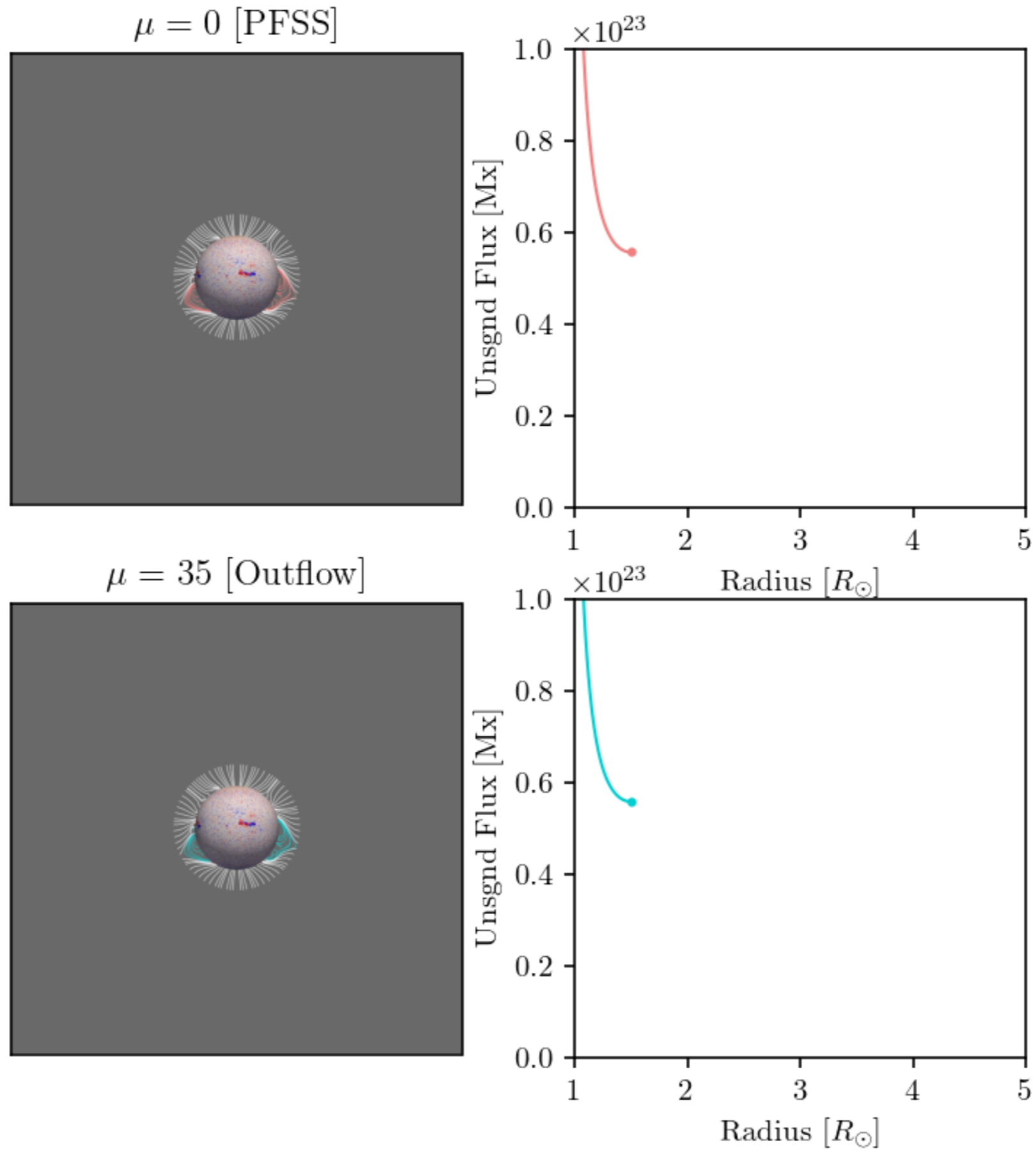
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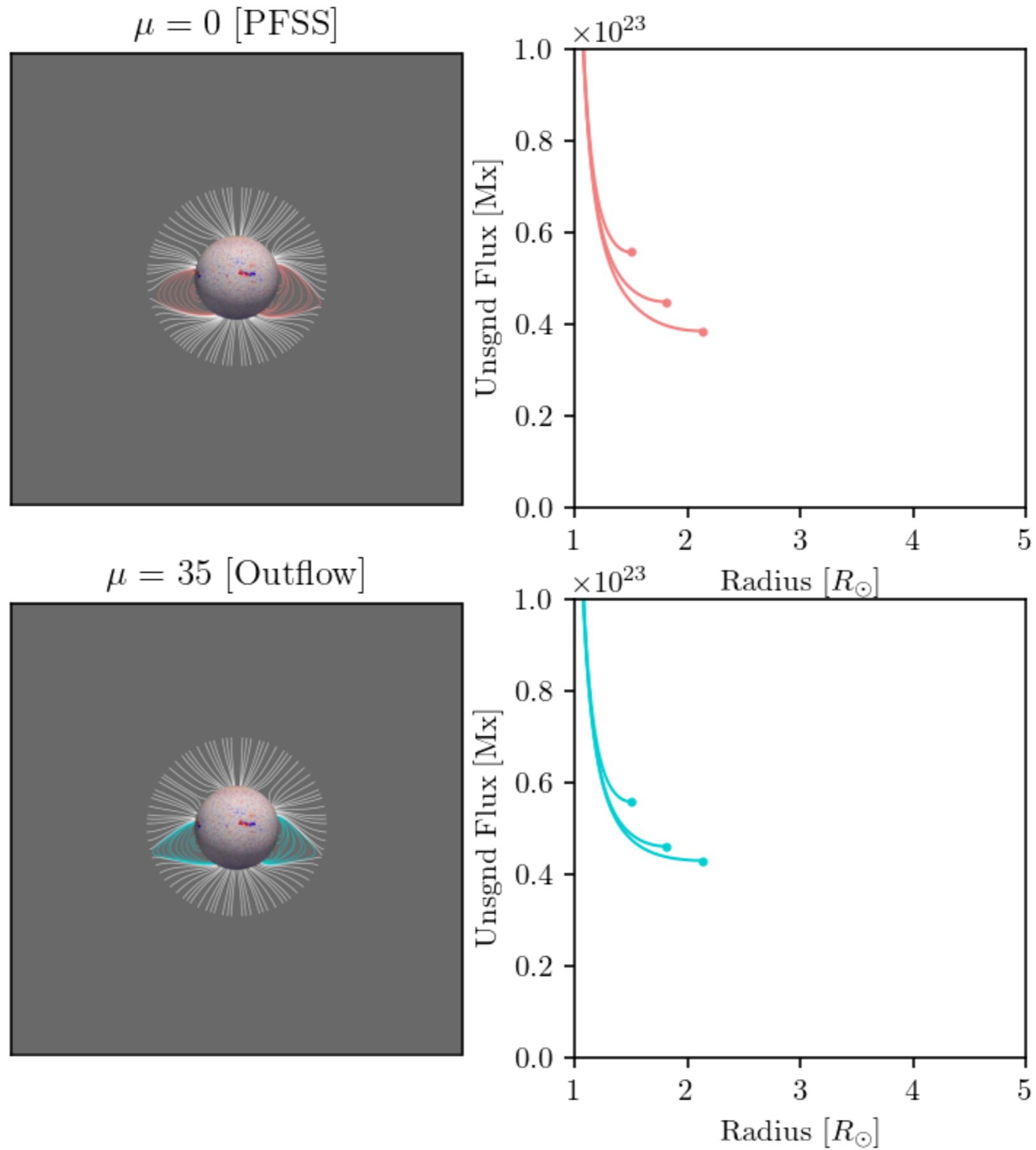
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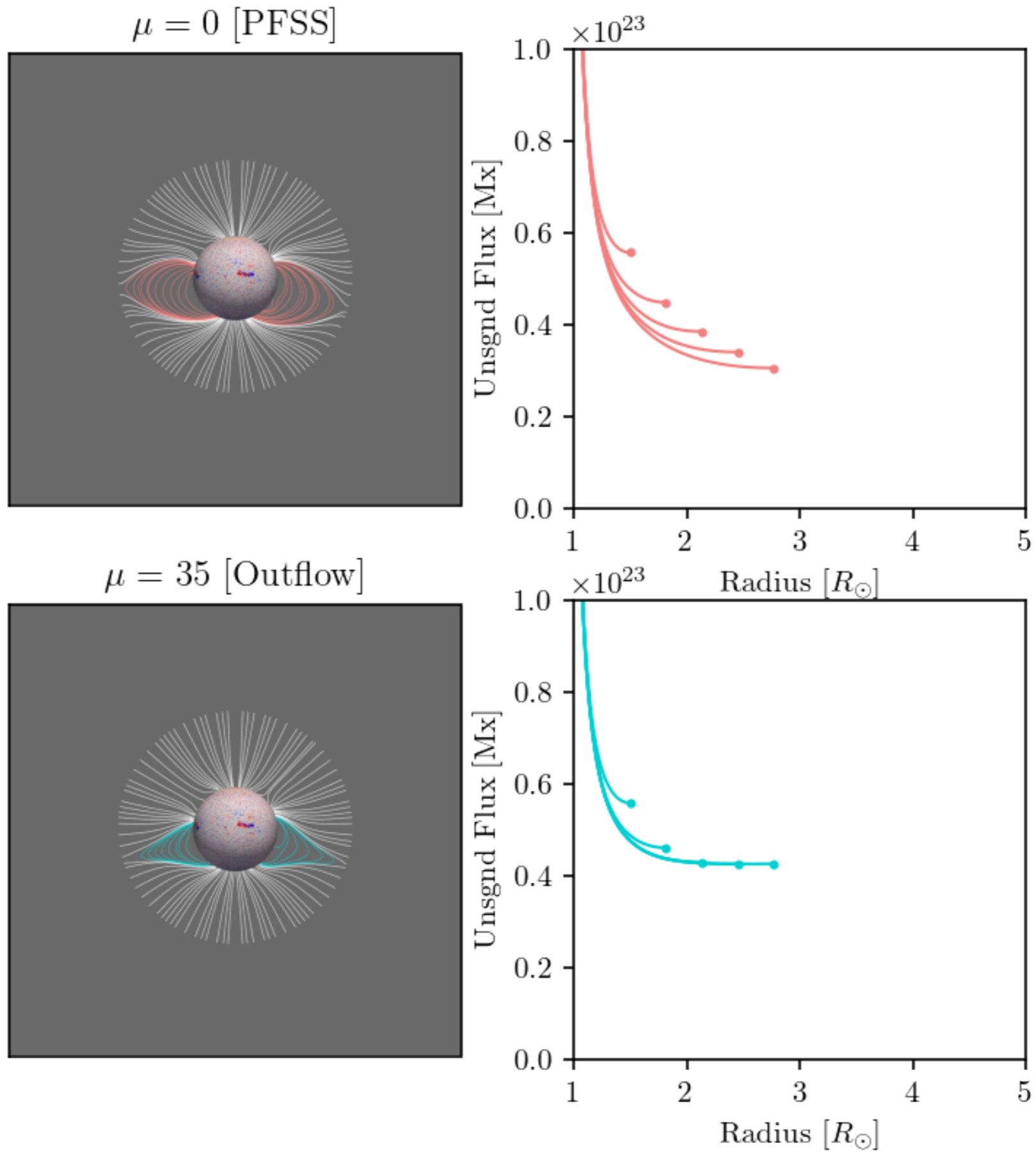
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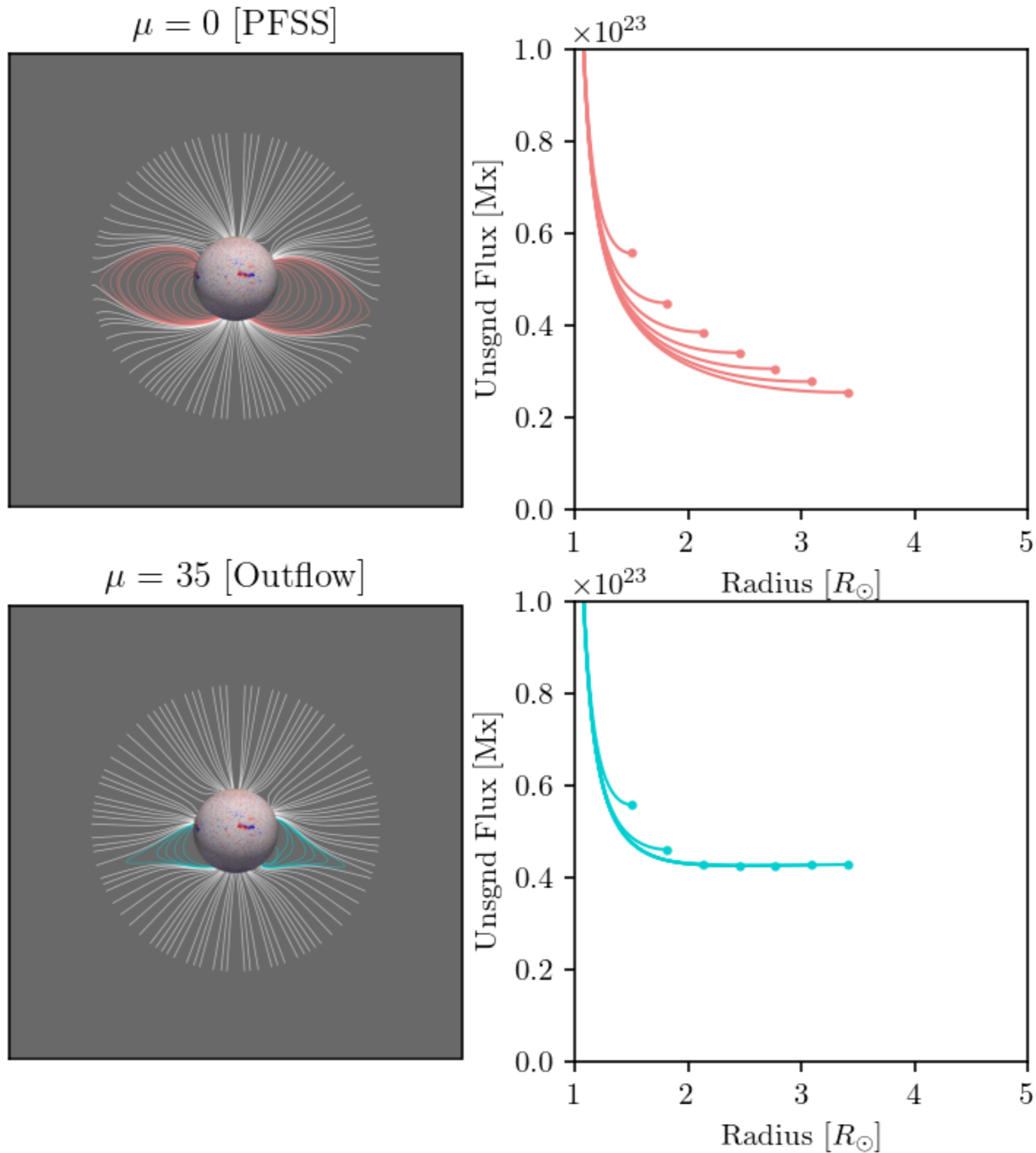
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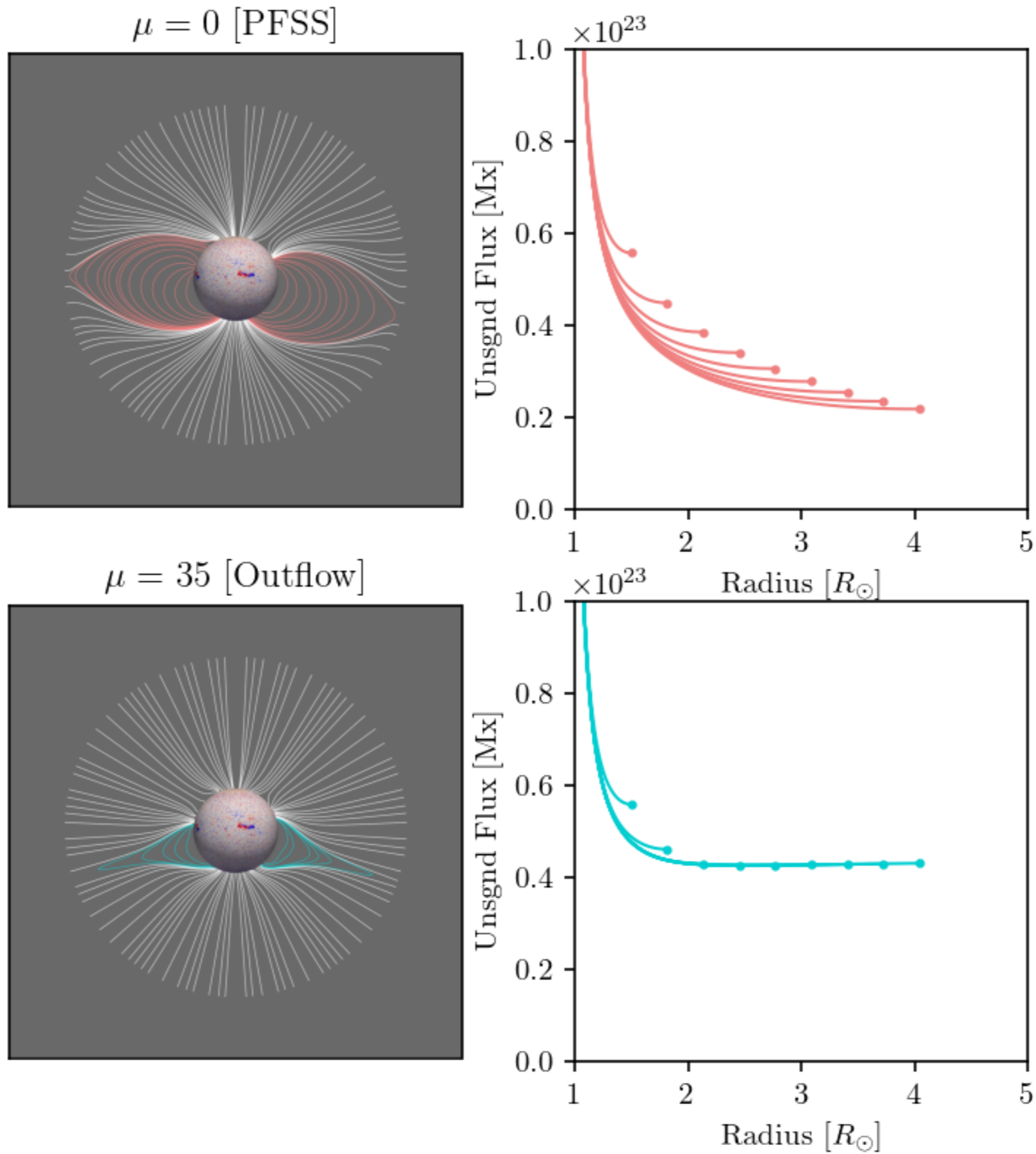
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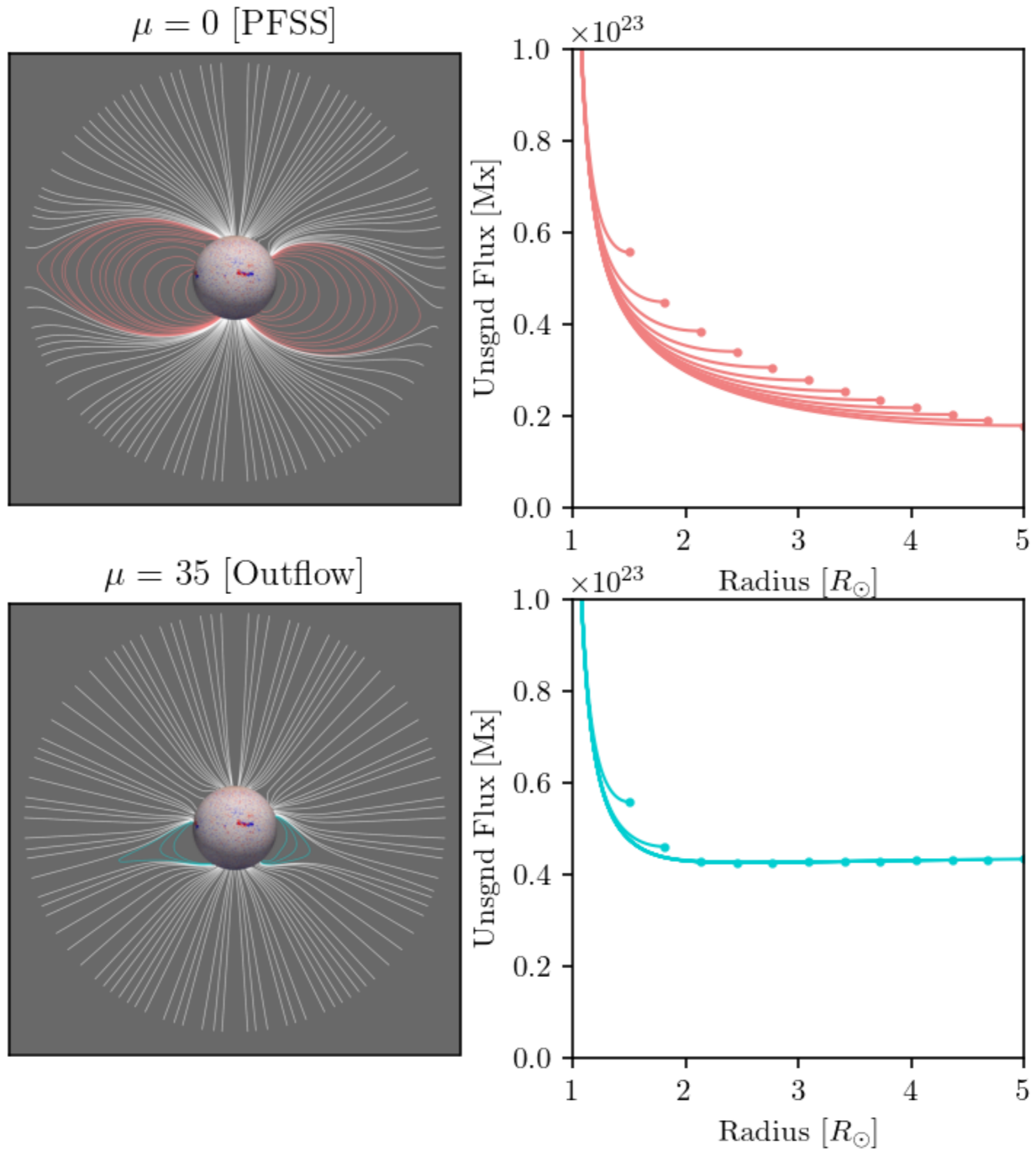
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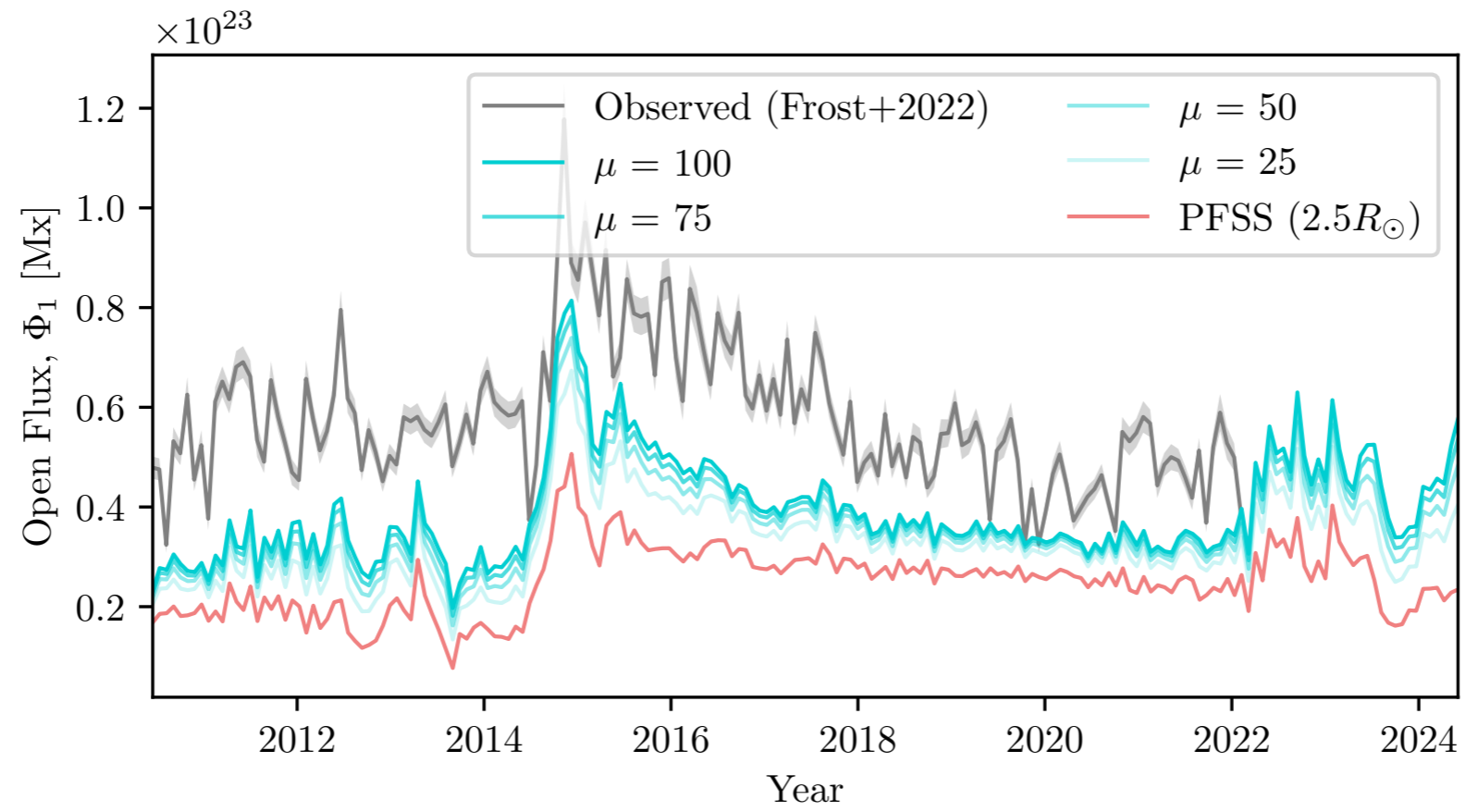


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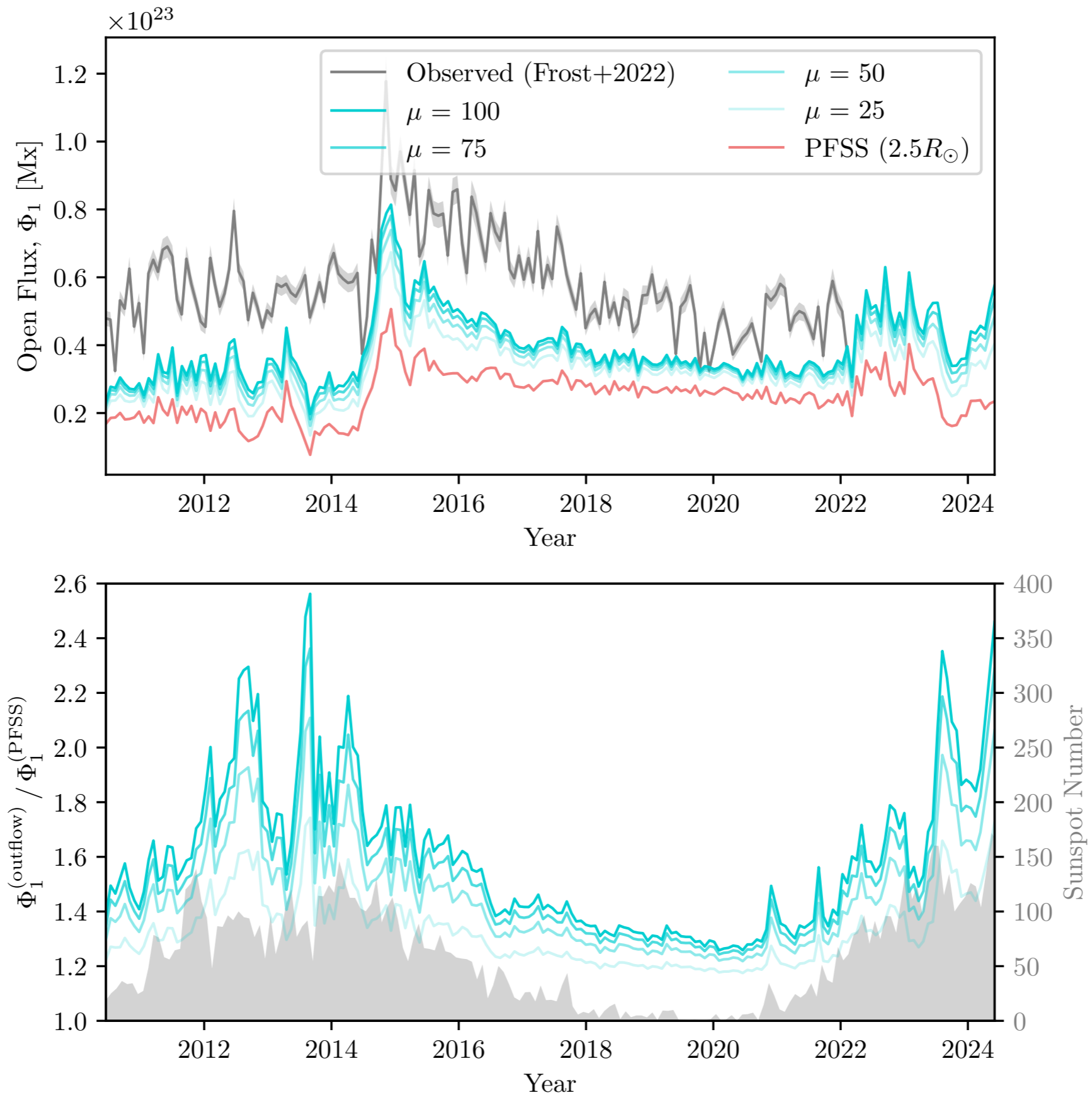


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Mathematical details...

Rice & Yeates, *ApJ* **932**, 57 (2021)

- ▶ Look for “magneto-frictional” equilibria of the form

$$\mathbf{v} \times \mathbf{B} = \mathbf{0}$$
$$\mathbf{v} - v_{\text{out}}(r)\hat{\mathbf{r}} = \frac{\mathbf{J} \times \mathbf{B}}{\nu_0 B^2}$$

$$\mathbf{J} = \nabla \times \mathbf{B}$$

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The diagram consists of a teal rectangular box on the left containing two equations. A teal arrow points from the right side of the box to the equation $\mathbf{B} = f(r)\nabla\psi$ on the right. Below the box is the equation $\mathbf{J} = \nabla \times \mathbf{B}$.

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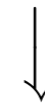
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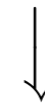
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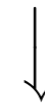
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- ▶ Depends on the dimensionless “Reynolds number” $\mu = R_{\odot}\nu_0 v_{\text{out}}(r_0)$